

Learning goals:

- Explain what is meant by the parallax of a star, how we measure it and use it to find the distance to a star.
- Define brightness, parallax effect, arc second, parsec, apparent magnitude, absolute magnitude.
- Describe the methods used to determine: temperature, luminosity, radius

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Questions:

Which stars are the brightest?

Which stars are putting out the most watts? (energy per second)

NEED TO KNOW:

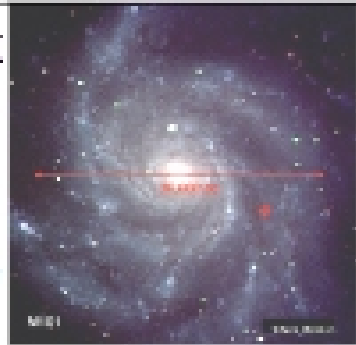
Distances

The most fundamental and accurate (within a certain range) means of finding distances is measuring the parallaxes of stars.

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Measured Parallax of Stars

- Works accurately for stars within about 200 pc (Hipparcos satellite)
- Biggest problem: measuring the miniscule shift of a star against more distant stars



parallax = 0.75 arcseconds

$$\text{distance} = \frac{1}{0.75} = 1.3 \text{ pc} = 4.3 \text{ ly}$$

parallax = 0.15 arcseconds

$$\text{distance} = \frac{1}{0.15} = 6.7 \text{ pc} = 22 \text{ ly}$$

parallax = 0.0015 arcseconds

$$\text{distance} = \frac{1}{0.0015} = 667 \text{ pc} = 2170 \text{ ly}$$

Define brightness, parallax effect, arc second, parsec, apparent magnitude, absolute magnitude

➤ Define parallax

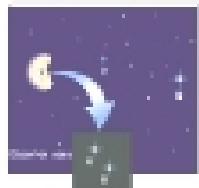
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Brightness

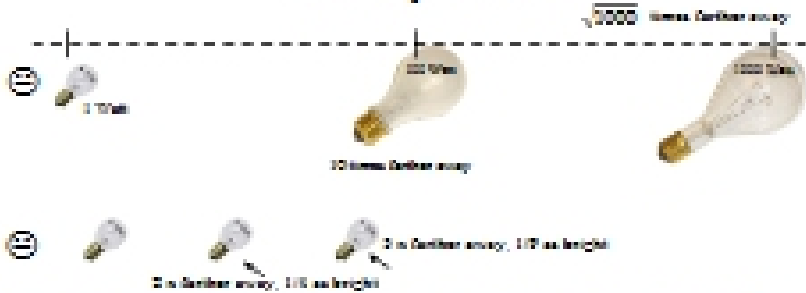
How the star looks to US HERE ON EARTH.

Depends On:

- Distance
- Luminosity

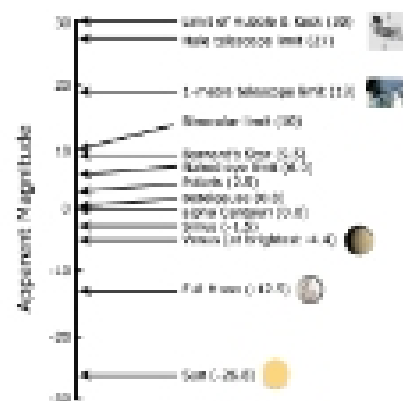


Inverse square law:



Define brightness, parallax effect, arc second, parsec, apparent magnitude, absolute magnitude

Apparent Magnitude



Every 5 magnitudes difference means 100 x difference in brightness

One magnitude difference is 2.512 times in brightness. (2.512⁵ = 100)

Define brightness, parallax effect, arc second, parsec, apparent magnitude, absolute magnitude

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